

Spatial structure of the modified Coulomb potential in a superstrong magnetic field

S. I. Glazyrin^{1,2}, S. I. Godunov²

¹Dukhov Research Institute of Automatics (VNIIA), Moscow, Russia

²Institute for Theoretical and Experimental Physics, NRC KI, Moscow, Russia
glazyrin@itep.ru

Ginzburg Centennial Conference on Physics
(May 29 – June 3, 2017)

Influence of magnetic field

- $a_B \sim a_H: B_1 = \frac{m^2 e^3 c}{\hbar^3} \approx 2.4 \times 10^9 \text{ G}$

$$U(r) = -\frac{e^2}{r} + \dots$$

- $\hbar\omega_B \sim mc^2: B_2 = \frac{m^2 c^3}{\hbar e} \approx 4.4 \times 10^{13} \text{ G}$

$$U(r) = -\frac{e^2}{r} + \dots$$

- $B_3 = \frac{m^2 c^4}{e^3} \approx 6 \times 10^{15} \text{ G}$
 $U(r)$ changes

Influence of magnetic field

- $a_B \sim a_H: B_1 = \frac{m^2 e^3 c}{\hbar^3} \approx 2.4 \times 10^9$ G

$$U(r) = -\frac{e^2}{r} + \dots$$

- $\hbar\omega_B \sim mc^2: B_2 = \frac{m^2 c^3}{\hbar e} \approx 4.4 \times 10^{13}$ G = B_0

Schwinger field

$$U(r) = -\frac{e^2}{r} + \dots$$

- $B_3 = \frac{m^2 c^4}{e^3} \approx 6 \times 10^{15}$ G – superstrong magnetic field
 $U(r)$ changes $U(r) = ?$

The modification of Coulomb potential

Coulomb potential is modified due to the enhancement of the vacuum polarization at one loop:

$$\Phi(\rho, z) = 4\pi e \int \frac{d^2 k_{\perp} dk_{\parallel}}{(2\pi)^3} \frac{e^{-i\vec{k}_{\perp}\vec{\rho}} e^{-ik_{\parallel}z}}{k_{\parallel}^2 + k_{\perp}^2 - \Pi^{(2)}(k_{\perp}, k_{\parallel})}$$

Polarization operator:

$$\Pi^{(2)}(k_{\perp}, k_{\parallel}) = -\frac{2e^3 B}{\pi} \exp\left(-\frac{k_{\perp}^2}{2eB}\right) T(t),$$

$$T(t) = 1 - \frac{1}{\sqrt{t(1+t)}} \log\left(\sqrt{1+t} + \sqrt{t}\right), \quad t \equiv k_{\parallel}^2/4m^2.$$

- One-loop calculation
- Lowest Landau Level

$$\begin{aligned}\hbar &= c = 1 \\ e^2 &= \alpha = 1/137.0..\end{aligned}$$

Numerical evaluation

Coulomb

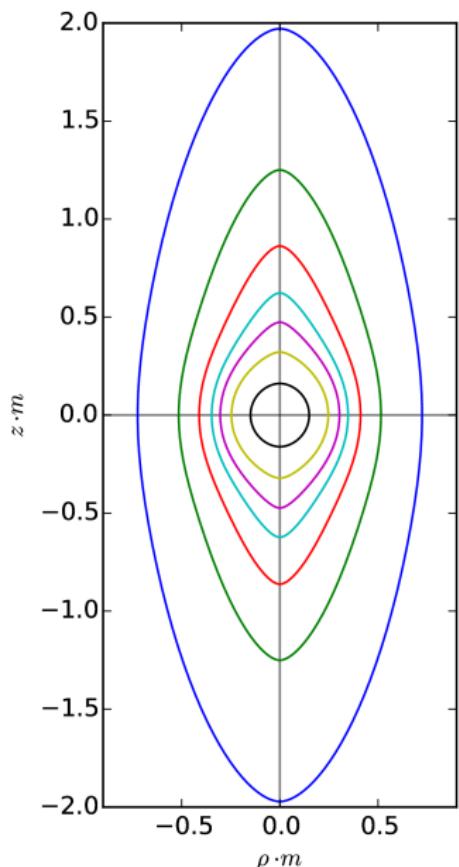
$$\begin{aligned}\Delta\Phi(\rho, z) &\equiv \frac{e}{\sqrt{\rho^2 + z^2}} - \Phi(\rho, z) = \\ &= \frac{e}{\pi} \int_{-\infty}^{\infty} dk_{\parallel} e^{-ik_{\parallel}z} \int_0^{\infty} dk_{\perp} k_{\perp} J_0(k_{\perp}\rho) \times \\ &\quad \times \frac{\frac{2e^3B}{\pi} e^{-k_{\perp}^2/2eB} T(k_{\parallel}^2/4m)}{\left(k_{\perp}^2 + k_{\parallel}^2\right) \left(k_{\perp}^2 + k_{\parallel}^2 + \frac{2e^3B}{\pi} e^{-k_{\perp}^2/2eB} T(k_{\parallel}^2/4m)\right)}.\end{aligned}$$

Two-step integration:

- ① Integration over k_{\perp} : Gnu Scientific Library (GSL)
- ② Integration over k_{\parallel} : Fast Fourier Transformation (FFTW)

The absolute error for any (ρ, z) is $< 10^{-6}(m \cdot e)$.

Potential structure for $B = 10^4 B_0$



Analytics:

- Small distances: Yukawa screening

$$\Phi(\rho, z) = \frac{e}{r} \cdot e^{-r\sqrt{2e^3B/\pi}},$$

$$r = \sqrt{\rho^2 + z^2}$$

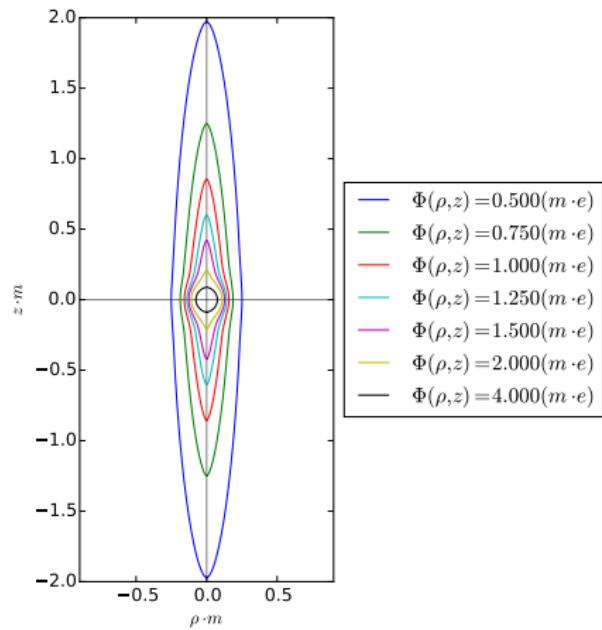
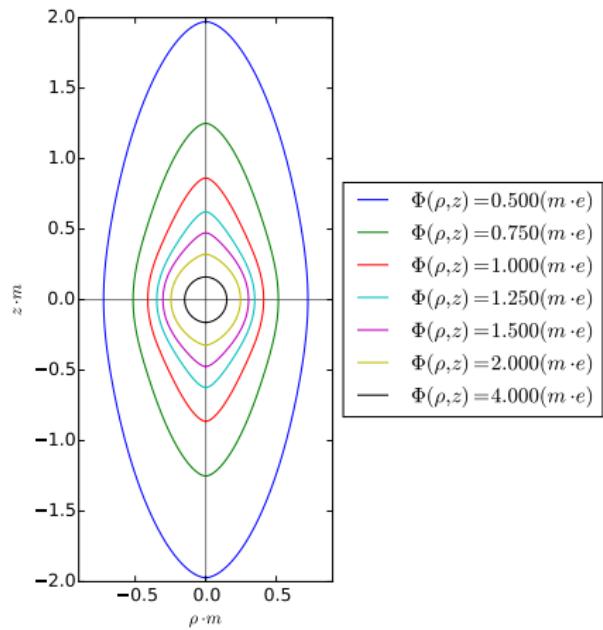
- Large distances: ellipses

$$\Phi(\rho, z) = \frac{e}{\sqrt{z^2 + \rho^2 \left(1 + \frac{e^3 B}{3\pi m^2}\right)}}$$

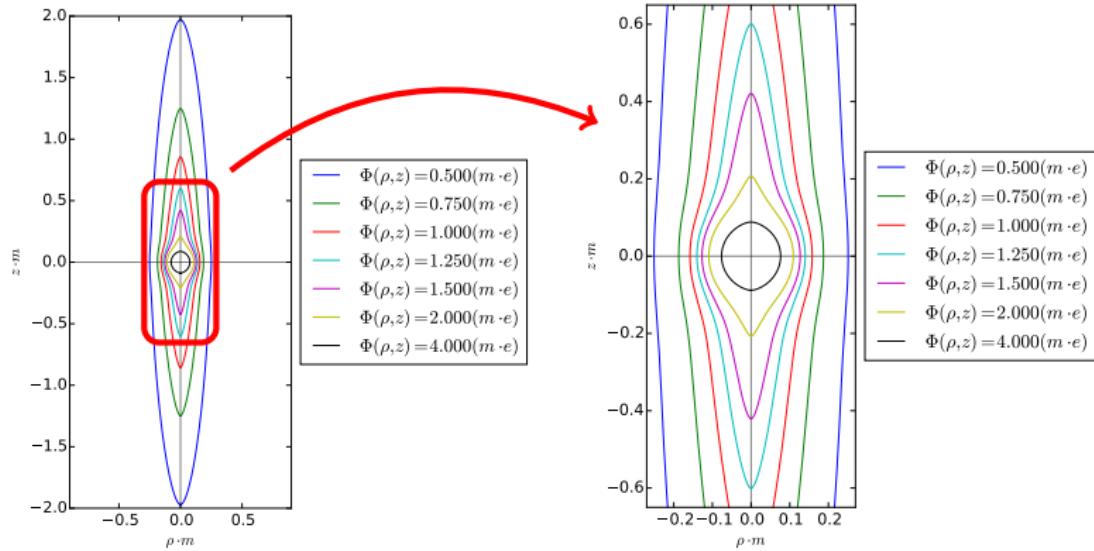
| | |
|---|------------------------------------|
| — | $\Phi(\rho, z) = 0.500(m \cdot e)$ |
| — | $\Phi(\rho, z) = 0.750(m \cdot e)$ |
| — | $\Phi(\rho, z) = 1.000(m \cdot e)$ |
| — | $\Phi(\rho, z) = 1.250(m \cdot e)$ |
| — | $\Phi(\rho, z) = 1.500(m \cdot e)$ |
| — | $\Phi(\rho, z) = 2.000(m \cdot e)$ |
| — | $\Phi(\rho, z) = 4.000(m \cdot e)$ |

Shabad, Usov (2007), (2008)
Vysotsky (2010)
Machet, Vysotsky (2011)
Godunov, Machet, Vysotsky
(2012)

Potential structure: $B = 10^4 B_0$ vs $B = 10^5 B_0$



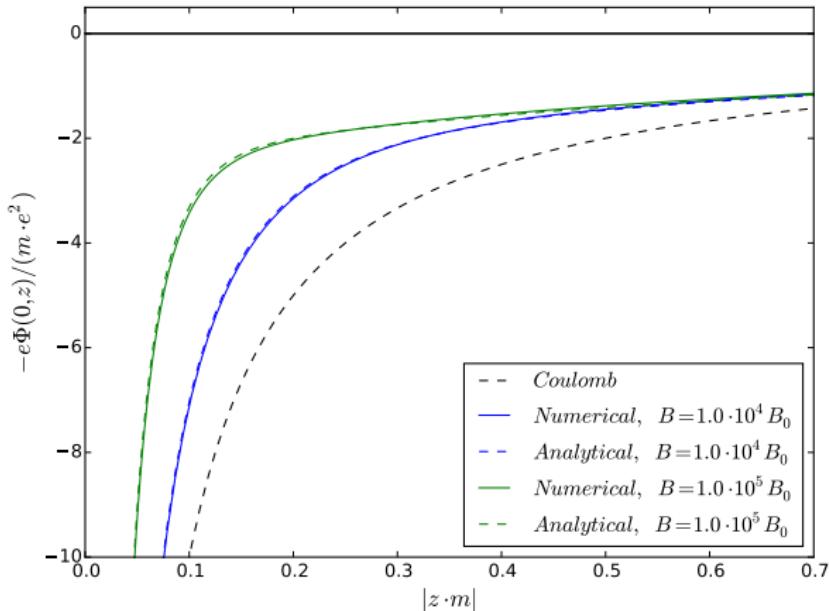
Internal structure for $B = 10^5 B_0$



Equipotential lines for $1/\sqrt{e^3 B} \lesssim z \lesssim 1/m$ are eye-shaped rather than ellipses

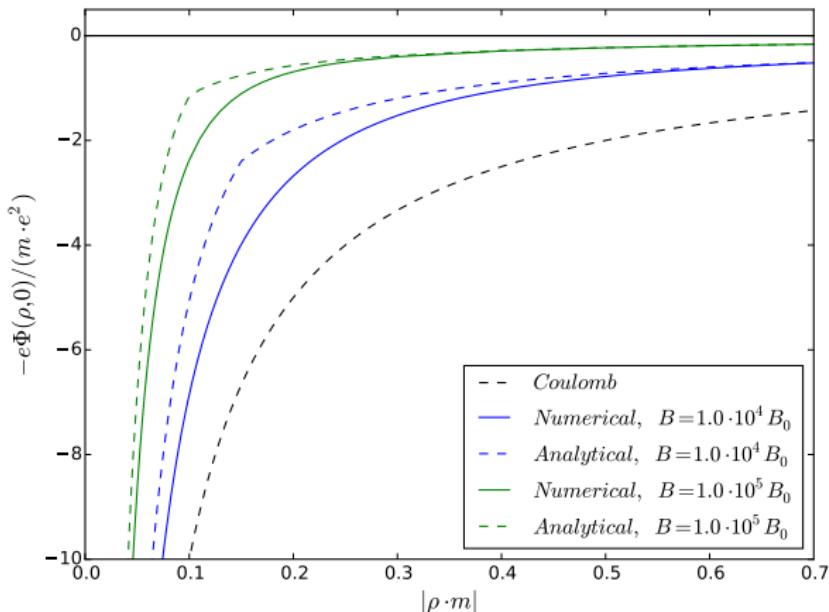
Potential for $(0, z)$

$$\Phi(0, z) = \frac{e}{|z|} \left(1 - e^{-|z|\sqrt{6m^2}} + e^{-|z|\sqrt{(2/\pi)e^3B+6m^2}} \right)$$



Potential for $(\rho, 0)$

$$\Phi(\rho, 0) = \begin{cases} \frac{e}{\rho} \cdot \exp\left(-\rho\sqrt{(2/\pi)e^3 B}\right), & \rho < l_0, \\ \frac{e}{\rho} \cdot \sqrt{\frac{3\pi m^2}{e^3 B}}, & \rho > l_0, \end{cases}$$



Possible applications

- Extreme magnetic fields in nature, e.g. for magnetars $B \sim 10^{15}$ G
- Condensed matter physics: $B_3 = \frac{m_{\text{eff}}^2}{e_{\text{eff}}^3}$.
- Critical charge problem
Oraevskii et al. (1977), Godunov et al. (2012), ...

Conclusions

- Superstrong magnetic field $B > m^2/e^3$ modifies the Coulomb potential (it becomes screened)
- We numerically calculated the modified potential in all space with high precision
- The simulations showed a new feature: at mid-range distances $1/\sqrt{e^3 B} \lesssim z \lesssim 1/m$ equipotential lines are eye-shaped
- Such a feature may be important for some problems, e.g. with spatially distributed charges (a direction for further study)

Thank you!